#### MODEL SELECTION

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## CURSE OF DIMENSIONALITY

- *k*-NN might yield much better results than a linear model, especially once we have chosen a good *k*.
- Maybe we won't need anything else?
- Let's see how *k*-NN behaves in high dimension (which is very realistic).

- Let us look for nearest neighbors for a point in a unit hypercube. Suppose that the original distribution was uniform.
- To cover share  $\alpha$  of test example, we have to cover (in expectation) a share  $\alpha$  of the volume, and the expected length of the side of a hypercube neighborhood in dimension p will be  $e_p(\alpha) = \alpha^{1/p}$ .
- E.g., in dimension  $10 e_{10}(0.1) = 0.8$ ,  $e_{10}(0.01) = 0.63$ , i.e., to cover 1% of the volume we have to take a neighborhood of length more than  $\frac{1}{2}$  w.r.t. each coordinate!
- This is bad for *k*-NN computationally too: it's hard to reject with a small number of coordinates, and fast algorithms don't work well.

- The second problem from the curse of dimensionality: consider *N* points uniformly distributed in a unit ball of dimension *p*.
- The mean distance to zero is

$$d(p,N) = \left(1 - \frac{1}{2}^{1/N}\right)^{1/p},$$

e.g., in dimension 10 for  $N = 500 \ d \approx 0.52$ , i.e., more than  $\frac{1}{2}$ .

 Most points are closer to the "sides" of the support than to other points, which is bad for k-NN: we extrapolate instead of interpolating.

- Third phenomenon: problems in optimization.
- To approximately optimize a function of *d* variables on a grid with step  $\epsilon$ , we will need approx.  $\left(\frac{1}{\epsilon}\right)^d$  function computations.
- Numerical integration: to integrate a function up to  $\epsilon$ , we will need  $\left(\frac{1}{\epsilon}\right)^d$  computations.

- Dense sets become very sparse. E.g., to get the density created in dimension 1 with N = 100 points we will need  $100^{10}$  points in dimension 10.
- The behaviour of functions also becomes more complicated as dimension grows: to construct regressions in high dimension with the same accuracy one might need exponentially more points than in low dimension.
- While a, say, linear model does not have any such effects, it's not subject to the curse of dimensionality.

• One more example: a normally distributed value will be concentrated in a thin shell.



**Exercise.** Convert the density of a Gaussian into polar coordinates and check this statement.

## EQUIVALENT KERNEL

• In linear regression we had

$$\begin{split} p(t \mid \mathbf{t}, \alpha, \beta) &= \mathcal{N}(t \mid \mu_N^\top \phi(\mathbf{x}), \sigma_N^2), \\ \text{where } \sigma_N^2 &= \frac{1}{\beta} + \phi(\mathbf{x})^\top \Sigma_N \phi(\mathbf{x}). \end{split}$$

• Let us rewrite the mean of the posterior as (recall that  $\mu_N = \beta \Sigma_N \Phi^\top \mathbf{t}$ ):

$$\begin{split} y(\mathbf{x}, \boldsymbol{\mu}_N) &= \boldsymbol{\mu}_N^\top \boldsymbol{\phi}(\mathbf{x}) = \boldsymbol{\beta} \boldsymbol{\phi}(\mathbf{x})^\top \boldsymbol{\Sigma}_N \boldsymbol{\Phi}^\top \mathbf{t} = \\ &= \sum_{n=1}^N \boldsymbol{\beta} \boldsymbol{\phi}(\mathbf{x})^\top \boldsymbol{\Sigma}_N \boldsymbol{\phi}(\mathbf{x}_n) t_n. \end{split}$$

- ·  $y(\mathbf{x},\boldsymbol{\mu}_N) = \sum_{n=1}^N \beta \phi(\mathbf{x})^\top \boldsymbol{\Sigma}_N \phi(\mathbf{x}_n) t_n.$
- And the prediction can be rewritten as

$$y(\mathbf{x},\boldsymbol{\mu}_N) = \sum_{n=1}^N k(\mathbf{x},\mathbf{x}_n) t_n.$$

- I.e., we predict the next point as a linear combination of values in known points.
- Function  $k(\mathbf{x}, \mathbf{x}') = \beta \phi(\mathbf{x})^\top \Sigma_N \phi(\mathbf{x}')$  is called the *equivalent* kernel.

#### EQUIVALENT KERNEL





- Equivalent kernel k(x, x') is localized around x as a function of x', i.e., every point has the largest influence nearby and then less and less (but it's not monotone!).
- We could simply define the kernel from the outset and predict with it without any  $\phi$  functions this is an important idea for the future.

**Exercise.** Prove that  $\sum_{n=1}^{N} k(\mathbf{x}, \mathbf{x}_n) = 1$ .

## **BAYESIAN MODEL COMPARISON**

- As the number of parameters increases, overfitting begins.
- How do we choose a model without overfitting? How can we compare models with different number of parameters?
- Bayesian approach: let's just compare  $p(\mathcal{M} \mid D)$ . :)

- Suppose we have a set of models  $\{\mathcal{M}_i\}_{i=1}^L$ .
- A model is a probability distribution over D.
- $\cdot\,$  And we can estimate the posterior

 $p(\mathcal{M}_i \mid D) \propto p(\mathcal{M}_i) p(D \mid \mathcal{M}_i).$ 

• If we know the posterior, we can make a prediction:

$$p(t \mid \mathbf{x}, D) = \sum_{i=1}^{L} p(t \mid \mathbf{x}, \mathcal{M}_i, \mathcal{D}) p(\mathcal{M}_i \mid D).$$

• *Model selection* is when we approximate the prediction by choosing the most probable model (a posteriori).

 $\cdot\,$  If the models are defined parametrically with  ${\bf w},$  we have

$$p(D \mid \mathcal{M}_i) = \int p(D \mid \mathbf{w}, \mathcal{M}_i) p(\mathbf{w} \mid \mathcal{M}_i) d\mathbf{w}.$$

- This is the probability to generate *D* if we choose model parameters with its prior and then sample the data.
- Exactly the denominator from the Bayes' theorem:

$$p(\mathbf{w} \mid \mathcal{M}_i, D) = \frac{p(D \mid \mathbf{w}, \mathcal{M}_i) p(\mathbf{w} \mid \mathcal{M}_i)}{p(D \mid \mathcal{M}_i)}$$

- Suppose that the model has a single parameter w, and the posterior is a sharp peak around  $w_{MAP}$  of width  $\Delta w_{posterior}$ .
- Then we can approximate  $p(D) = \int p(D \mid w)p(w)dw$  as the value in the maximum times the width.
- Let's also assume that the prior distribution is flat,

 $p(w) = \frac{1}{\Delta w_{\text{prior}}}.$ 

# Approximating $p(\boldsymbol{d})$



### Approximating p(d)

 $\cdot$  Then we get

$$\begin{split} p(D) &= \int p(D \mid w) p(w) dw \approx p\left(D \mid w_{\mathrm{MAP}}\right) \frac{\Delta w_{\mathrm{posterior}}}{\Delta w_{\mathrm{prior}}}, \\ &\ln p(D) \approx \ln p\left(D \mid w_{\mathrm{MAP}}\right) + \ln \left(\frac{\Delta w_{\mathrm{posterior}}}{\Delta w_{\mathrm{prior}}}\right). \end{split}$$

- This means that we add a penalty for "too narrow" posteriors... that is, precisely the penalty for overfitting!
- + For a model of M parameters, if we assume that they have identical  $\Delta w_{\rm posterior}$  we get

$$\ln p(D) \approx \ln p\left(D \mid w_{\text{MAP}}\right) + M \ln \left(\frac{\Delta w_{\text{posterior}}}{\Delta w_{\text{prior}}}\right).$$

- In other words: let's see what kinds of datasets can be generated by a certain model.
- A simple model (e.g., linear) generates similar datasets, "few" different datasets, and so has high  $p(D \mid M)$ .
- A complicated model (e.g., degree 9 poly) generates "many" different datasets, and so has low  $p(D \mid \mathcal{M})$ .
- But a complicated model can express datasets that a simple one cannot; so in total we should choose a "middle ground"

# Approximating $p(\boldsymbol{d})$



- Sanity check: we have introduced strange-looking penalties; but will the true correct answer  $p(D \mid \mathcal{M}_{true})$  be actually optimal in this sense?
- For a specific dataset, not necessarily.
- But averageing over all datasets sampled from the true distribution  $p(D \mid \mathcal{M}_{\mathrm{true}}) ...$

• ...we get

$$\mathbf{E}\left[\ln\frac{p(D\mid\mathcal{M}_{\mathrm{true}})}{p(D\mid\mathcal{M})}\right] = \int p(D\mid\mathcal{M}_{\mathrm{true}})\ln\frac{p(D\mid\mathcal{M}_{\mathrm{true}})}{p(D\mid\mathcal{M})}dD.$$

• This is called Kullback-Leibler divergence between distributions  $p(D \mid \mathcal{M}_{true})$  and  $p(D \mid \mathcal{M})$ . Exercise. Prove that the Kullback-Leibler divergence is always nonnegative, i.e.,  $p(D \mid \mathcal{M}_{true}) \geq p(D \mid \mathcal{M})$  for every  $\mathcal{M}$ .

#### MODEL COMPARISON WITH LAPLACE APPROXIMATION

- But we can do better than just a flat plateau! Let's try Laplace approximation for this case.
- Again, to compare models  $\{\mathcal{M}_i\}_{i=1}^L$  , we evaluate with test set D the posterior

 $p(\mathcal{M}_i \mid D) \propto p(\mathcal{M}_i) p(D \mid \mathcal{M}_i).$ 

- If the models are parametric then  $p(D \mid \mathcal{M}_i) = \int p(D \mid \theta, \mathcal{M}_i) p(\theta \mid \mathcal{M}_i) d\theta.$
- This is the probability to generate *D* if we choose model parameters according to its prior, the Bayes theorem's denominator:

$$p(\theta \mid \mathcal{M}_i, D) = \frac{p(D \mid \theta, \mathcal{M}_i)p(\theta \mid \mathcal{M}_i)}{p(D \mid \mathcal{M}_i)}.$$

#### MODEL COMPARISON WITH LAPLACE APPROXIMATION

• Let's approximate with a Gaussian; integrating, we get

$$Z = \int f(\mathbf{z}) d\mathbf{z} \approx \int f(\mathbf{z}_0) e^{-\frac{1}{2} (\mathbf{z} - \mathbf{z}_0)^\top \mathbf{A}(\mathbf{z} - \mathbf{z}_0)} d\mathbf{z} = f(\mathbf{z}_0) \frac{(2\pi)^{M/2}}{|\mathbf{A}|^{1/2}}.$$

• And we have Z = p(D),  $f(\theta) = p(D \mid \theta)p(\theta)$ .

• We get

$$\ln p(D) \approx \ln p(D \mid \theta_{\text{MAP}}) + \ln P(\theta_{\text{MAP}}) + \frac{M}{2}\ln(2\pi) - \frac{1}{2}\ln|\mathbf{A}|.$$

- $\ln P(\theta_{\text{MAP}}) + \frac{M}{2}\ln(2\pi) \frac{1}{2}\ln|\mathbf{A}|$  is called the Occam's factor.
- $\boldsymbol{\cdot} \ \mathbf{A} = -\nabla \nabla \ln p(D \mid \boldsymbol{\theta}_{\mathrm{MAP}}) p(\boldsymbol{\theta}_{\mathrm{MAP}}) = -\nabla \nabla \ln p(\boldsymbol{\theta}_{\mathrm{MAP}} \mid D).$

• We get

$$\ln p(D) \approx \ln p(D \mid \theta_{\mathrm{MAP}}) + \ln P(\theta_{\mathrm{MAP}}) + \frac{M}{2}\ln(2\pi) - \frac{1}{2}\ln|\mathbf{A}|.$$

• If the Gaussian prior  $p(\theta)$  is sufficiently wide, and **A** has full rank, we can roughly approximate as (prove it!)

$$\ln p(D) \approx \ln p(D \mid \boldsymbol{\theta}_{\mathrm{MAP}}) - \frac{1}{2} M \ln N,$$

where M is the number of parameters, N is the number of points in D, and we have omitted additive constants.

• This is called the *Bayesian information criterion* (BIC), or *Schwarz criterion*.

# Thank you for your attention!